Brian Hicks

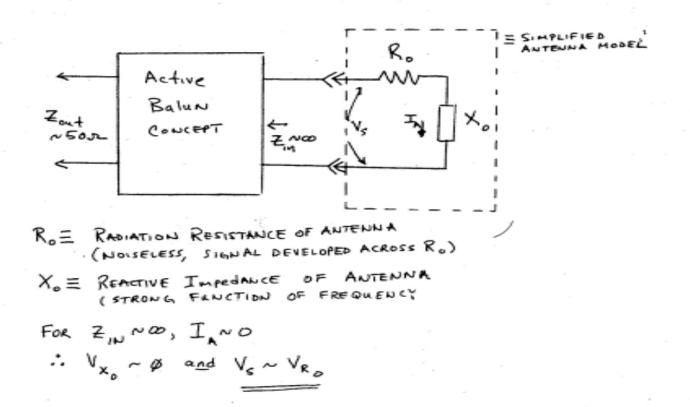
W. Erickson, K. Stewart

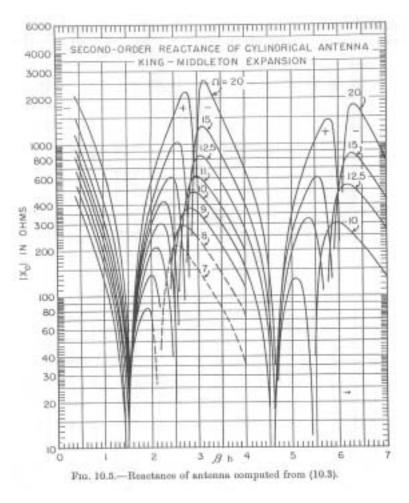
Our *goal* has been to research and develop widebandwidth low-frequency antennas to operate from 10 MHz (~ionospheric cut-off) to approximately 120 MHz for use with LOFAR.

Our *approach* has been to model, prototype, and test antenna designs by observing actual astronomical sources (e.g. Galactic drift scans, solar bursts, and other strong sources)

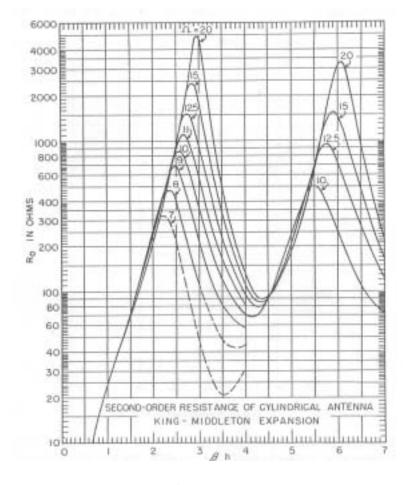
NRL Low Frequency Antenna Development Antenna System Components

- Active Baluns Early NRL efforts focused on using high input impedance amplifiers coupled directly to folded dipole antennas. It was our intent to mitigate the widely varying impedances associated with such antennas at lower frequencies.
- Antenna Elements Early NRL efforts focused on folded-dipole antennas.





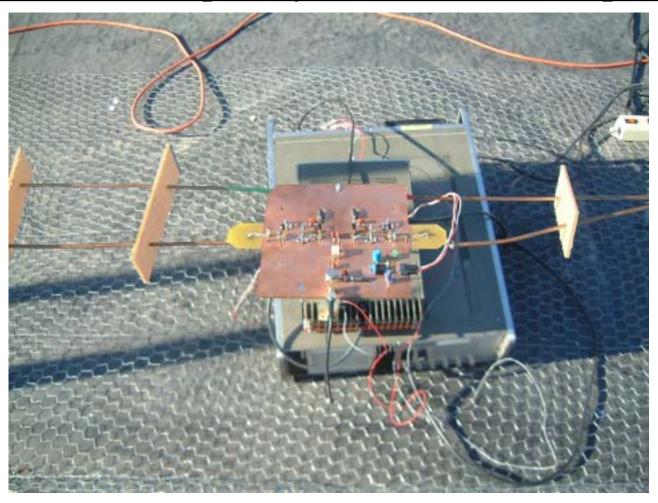
Antenna Reactance vs. Frequency



Antenna Resistance vs. Frequency



Active-dipole tests based on folded dipole



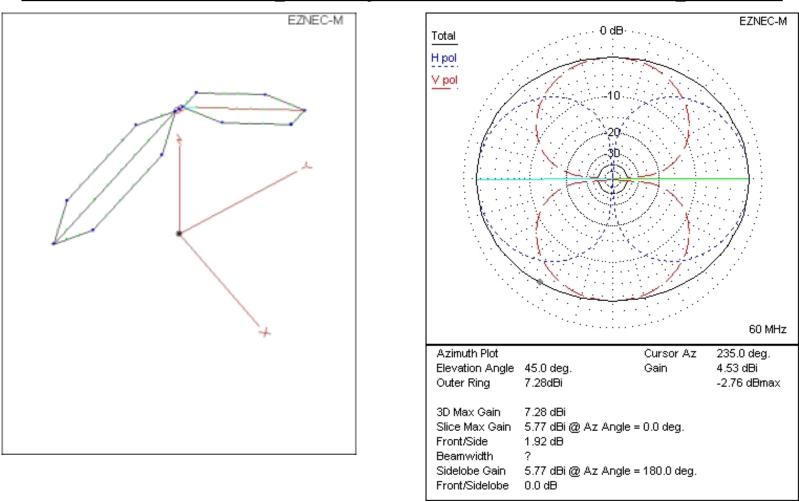
Early active-balun design coupled to folded-dipole



Active-balun prototype utilizing commercial wide-bandwidth Operational Amplifier



Newer designs focus on mainstream components such as CATV amplifiers

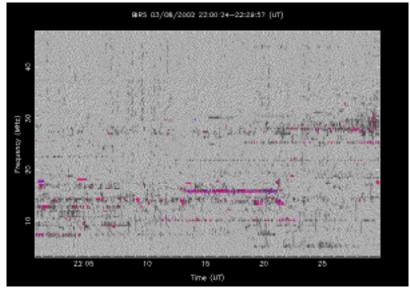


Design and Simulation of Inverted-V Dipole Antenna



Log-Periodic at BIRS Site (Bruny Island Radio Spectrometer)

William Erickson's BIRS system conducts solar burst observations in the 3 to 20 MHz range. In operation since 1995, BIRS utilizes digital avoidance and excision techniques to minimize terrestrial RFI. BIRS data is updated daily on its WWW site.



BIRS data from 3-8-2002

BIRS Web-Site: http://fourier.phys.utas.edu.au/birs/



Ken Stewart



William Erickson

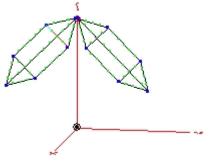
Groundbreaking for Recently Completed Field Tests at William Erickson's BIRS Site (Bruny Island, Tasmania)



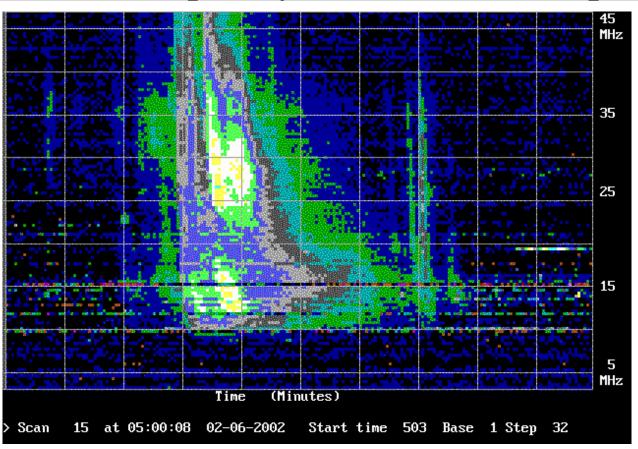
Ground Screen





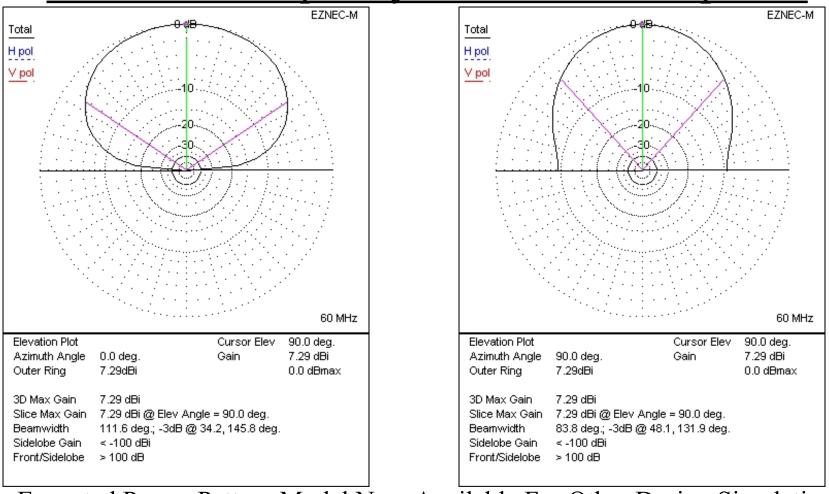


CAD Model



Type V Solar Burst Observed with Prototype LF Antenna

The prototype antenna was integrated into the BIRS system and astronomical observations (e.g. Galactic drift scans, solar burst observations) were conducted.



Expected Power Pattern Model Now Available For Other Design Simulations available to all LOFAR consortium members

Goals and Milestones

- Present until November, 2003:

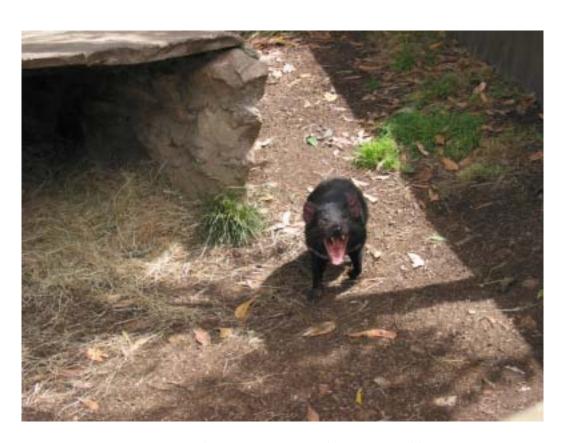
Form partnership with commercial antenna manufacturer to optimize design for mass production and manage risk.

- Present until July, 2002:

Develop two element interferometer test system to provide refined antenna performance data to LOFAR consortium.

- Present until November, 2003:

Develop prototype stations using cross-polarized LF antenna design.



Genuine Tasmanian Devil



Popular Anthropomorphization